



Program Options to Explore Ocean Worlds

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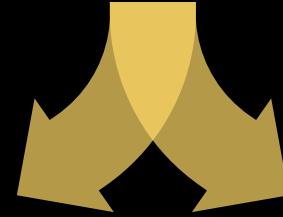
42nd COSPAR Assembly, Pasadena, 20 July 2018



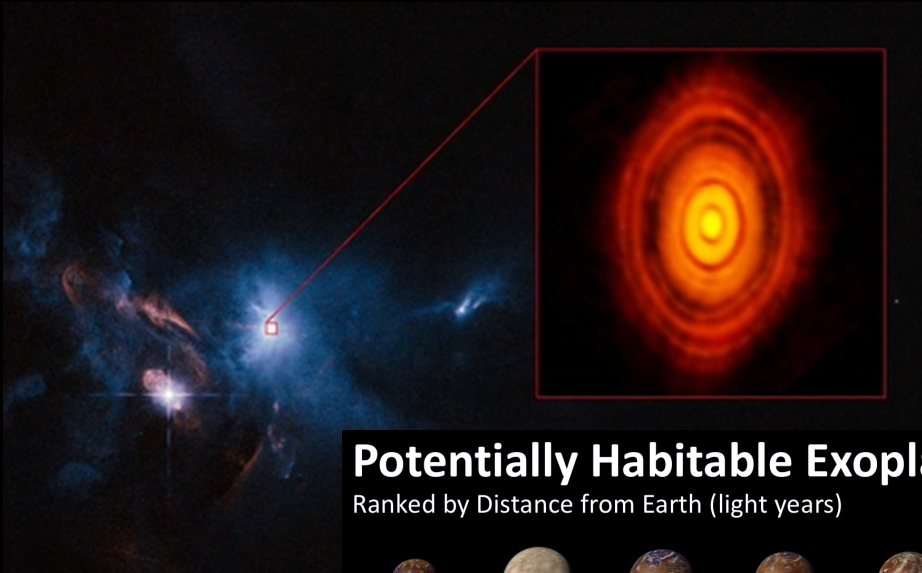
Jet Propulsion Laboratory
California Institute of Technology

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Is there life elsewhere in the universe?



Indirect search:
observation of
extra-solar planets

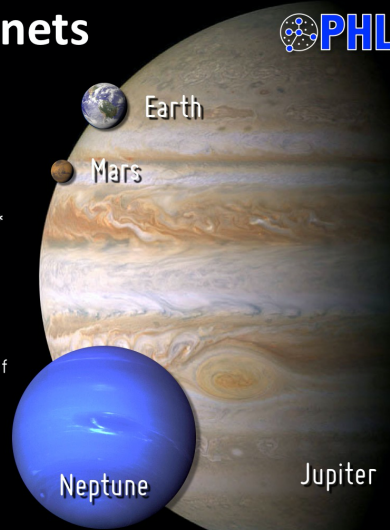


Potentially Habitable Exoplanets

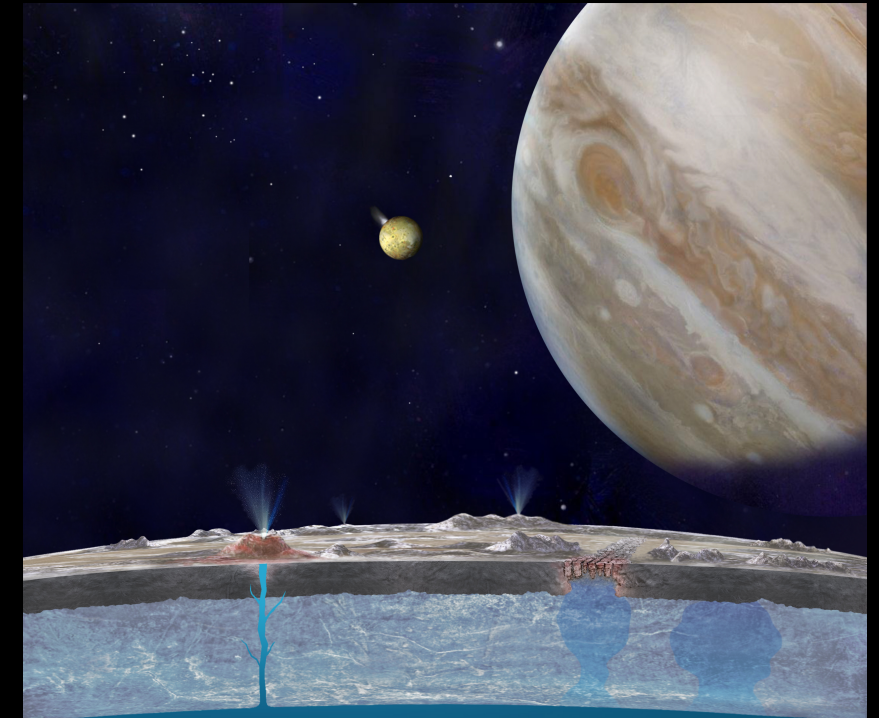
Ranked by Distance from Earth (light years)

 [4 ly] Proxima Cen b	 [13 ly] Kapteyn b*	 [22 ly] GJ 667 C c	 [22 ly] GJ 667 C f*	 [22 ly] GJ 667 C e*
 [39 ly] TRAPPIST-1 e	 [39 ly] TRAPPIST-1 f	 [39 ly] TRAPPIST-1 g	 [41 ly] LHS 1140 b	 [561 ly] Kepler-186 f
 [770 ly] Kepler-1229 b	 [822 ly] Kepler-1652 b	 [1115 ly] Kepler-442 b	 [1200 ly] Kepler-62 f	

Artistic representations. Earth, Mars, Jupiter, and Neptune for scale. Distance from Earth is between brackets. Planet candidates indicated with asterisks.



CREDIT: PHL @ UPR Arecibo (phl upr.edu) Jul 2, 2018



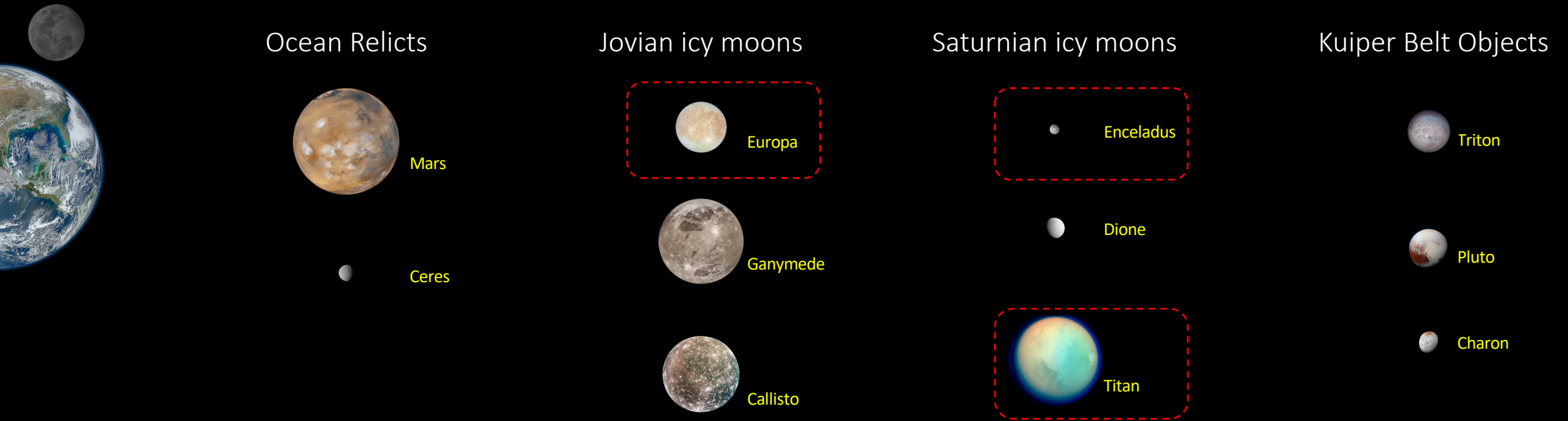
Tangible search:
life in the ocean worlds
of our solar system

Ocean worlds contain vast quantities of water



EARTH

More than a dozen ocean worlds within reach



Six steps to “find and understand life elsewhere”

1. Find liquid water
2. Quantify its habitability
3. Detect biosignatures in it
4. Confirm that life is present
5. Understand how that life operates
6. Learn the limits of life

Program scenarios for a NASA OWE

Case 0a – Status quo

Case 0b – Select Dragonfly in 2019

Case 1 or 2 would cost
 $1/40^{\text{th}}$ of NASA budget

Case 1 – Select two New Frontiers missions every 5 years, one for OW

Case 2 – Compete \$1B-class directed-purpose OW mission every 5 years

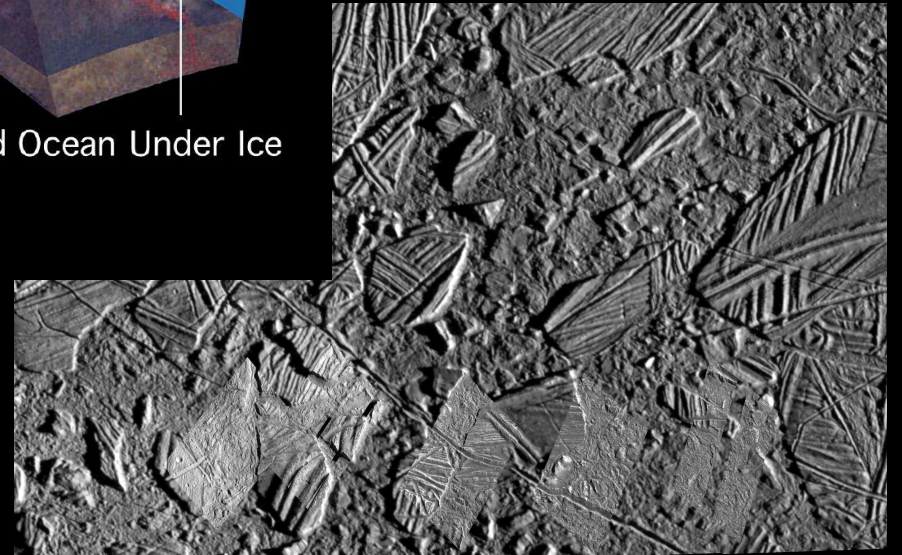
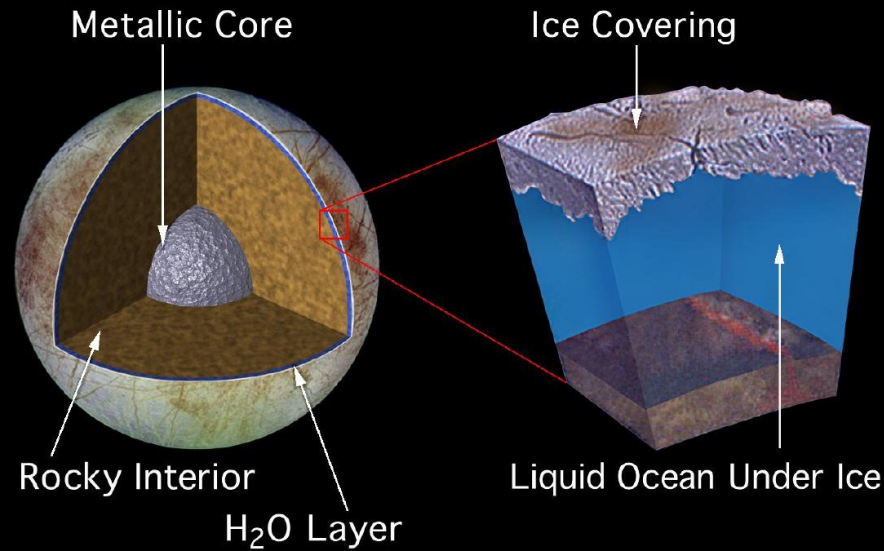
Case 3a – Add strategically managed OW Technology Program

Case 3b – Establish a line-item OW Exploration Program

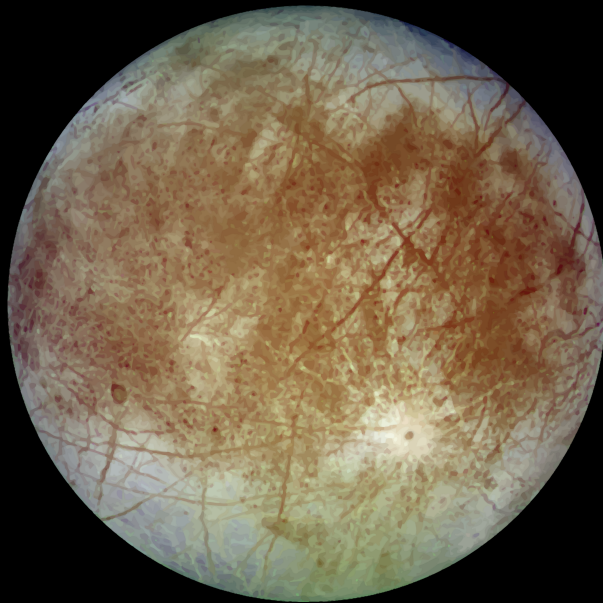
The cost information contained in this document is of a budgetary and planning nature and is intended for informational purposes only. It does not constitute a commitment on the part of JPL and/or Caltech.

Europa: best place for life?

- Almost as large as Earth's Moon
- Ocean in contact with silicate rock
- Plausible mechanism to sustain redox disequilibrium

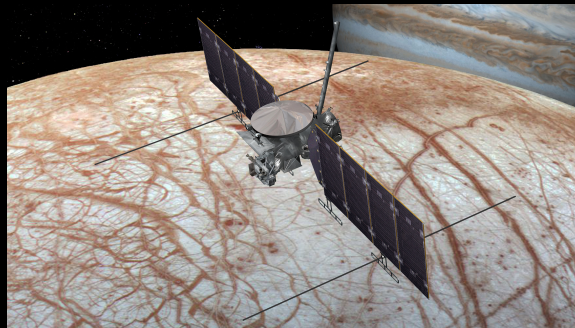


A Europa program in three steps



Comprehensive investigation of the icy moon's habitability

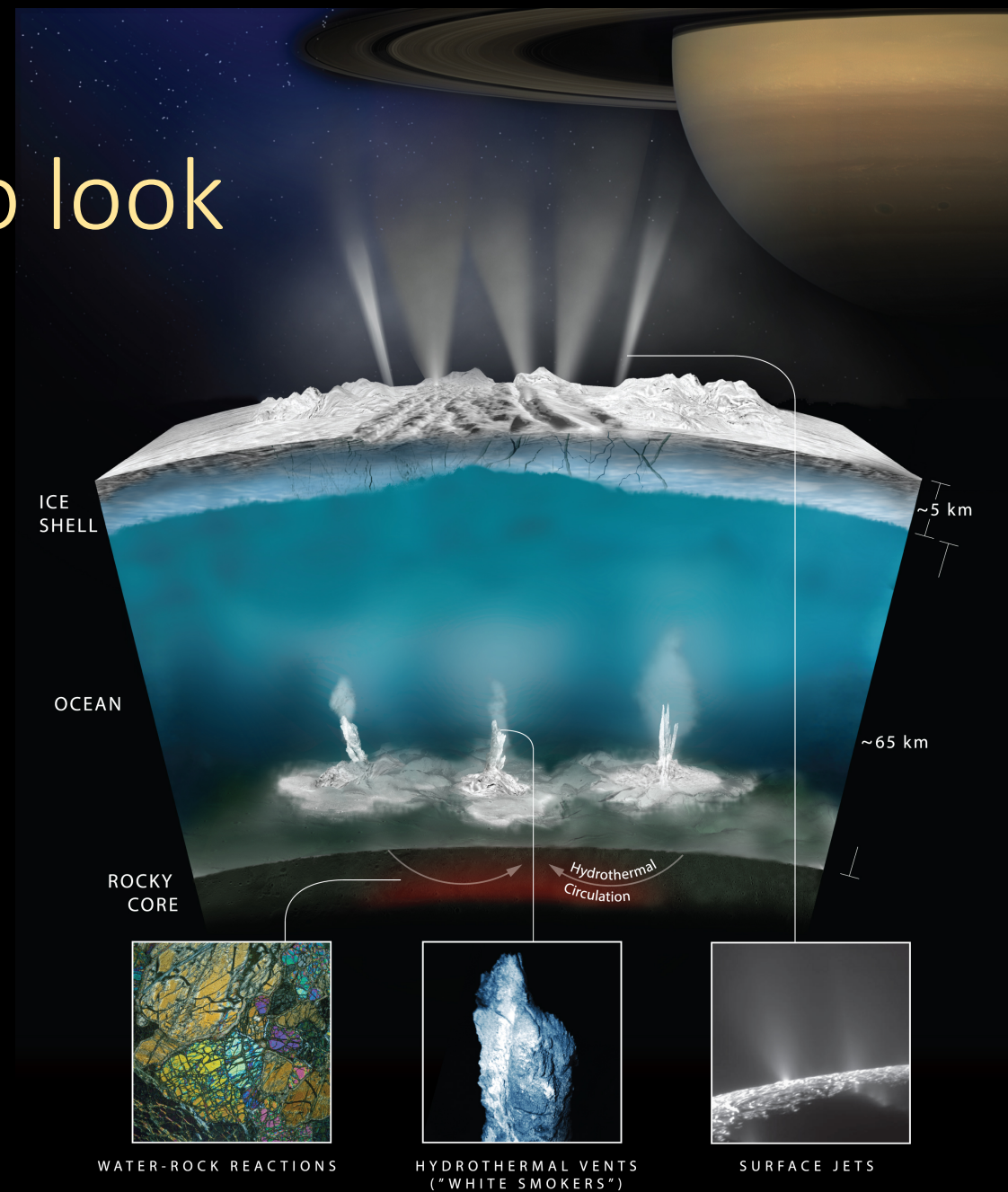
- Near-global hyper mapping
- Lander-scale surface imagery



- Land at ocean-surface exchange zone
- Mobility around touchdown point
- Subsurface access to pursue fresher material
- Trans-shell probe into ocean, sample return
- Under-ice exploration of ocean ceiling
- Open ocean exploration, including seafloor

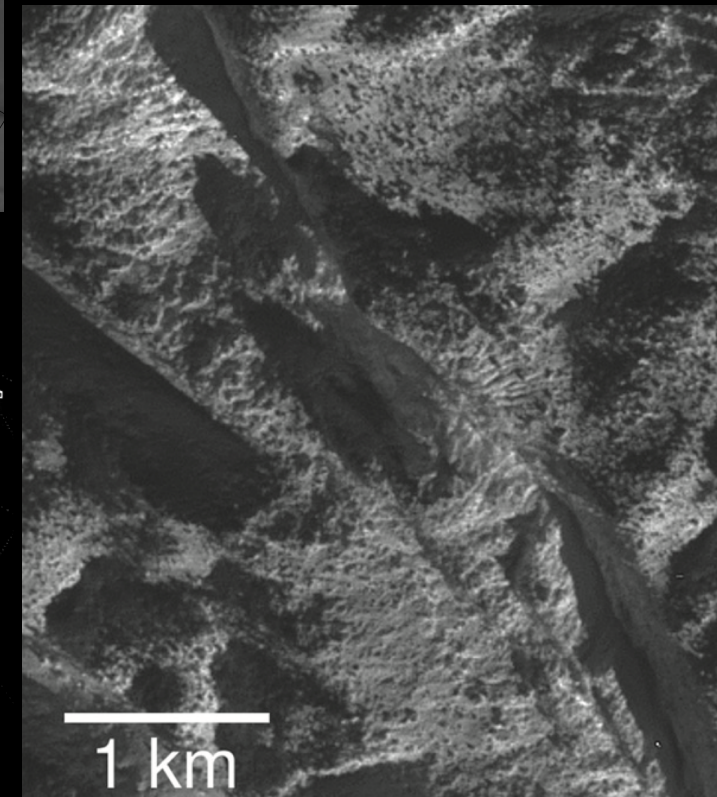
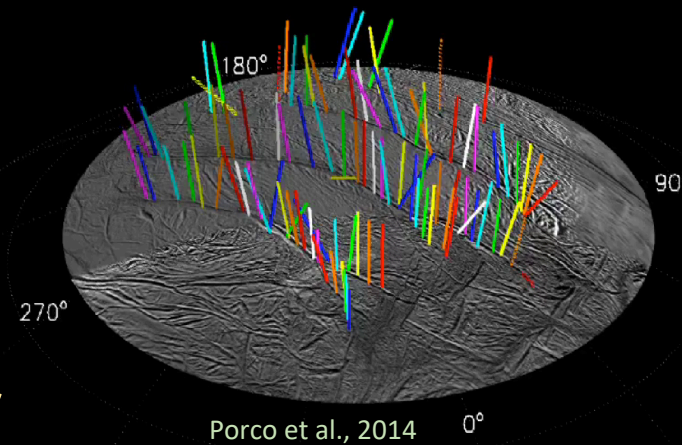
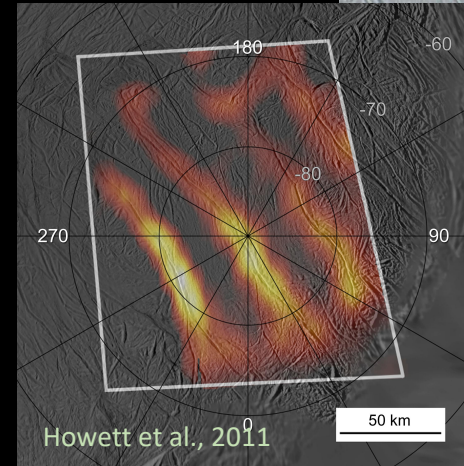
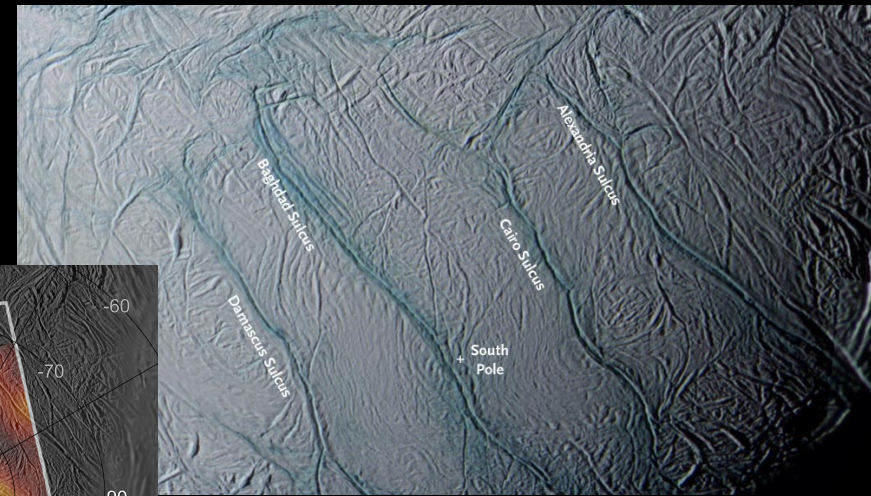
Enceladus: easiest place to look

- The most habitable place we know so far
- Salt-water ocean with hydrothermal activity
- Predictably expressed into space by a big plume, ripe for sampling

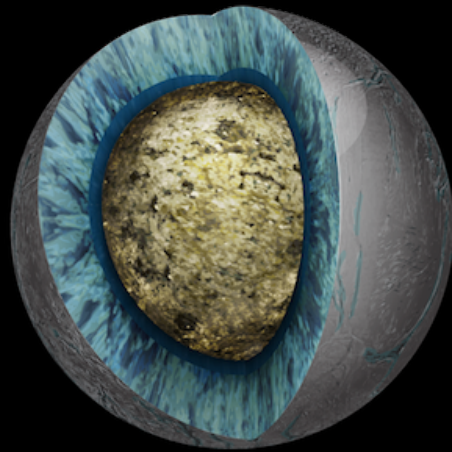


Awaiting focused missions

- Ice crust split by warm fissures
- > 100 geysers send large plume far into space, easy to transect
- Ice grains include silica dust and frozen ocean spray
- Salt water, hydrothermal circulation, organic chemistry



An Enceladus program in three steps



Direct access to material known to originate in a habitable place

- Plume transects
- Best compositional analyzers

Wet-chemistry and microscopy of grain material

Collection, preservation, and return of samples

Surface collection of large amounts of material

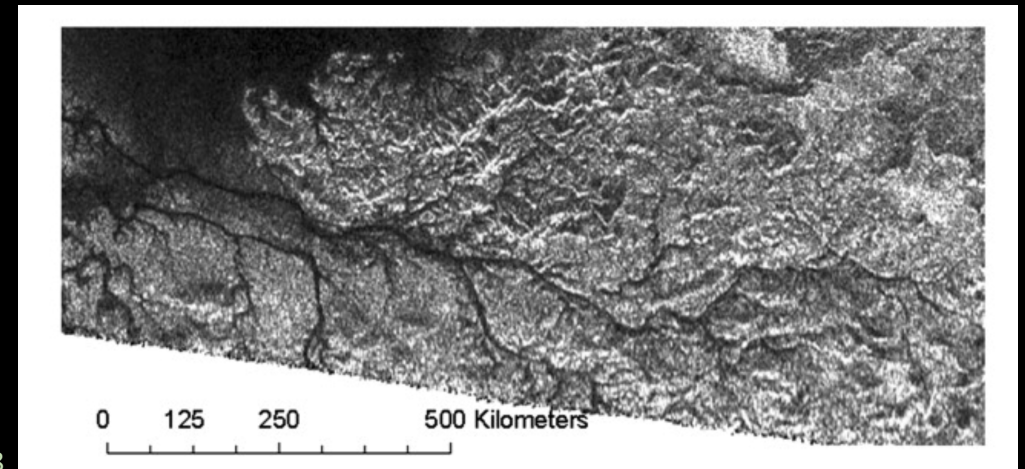
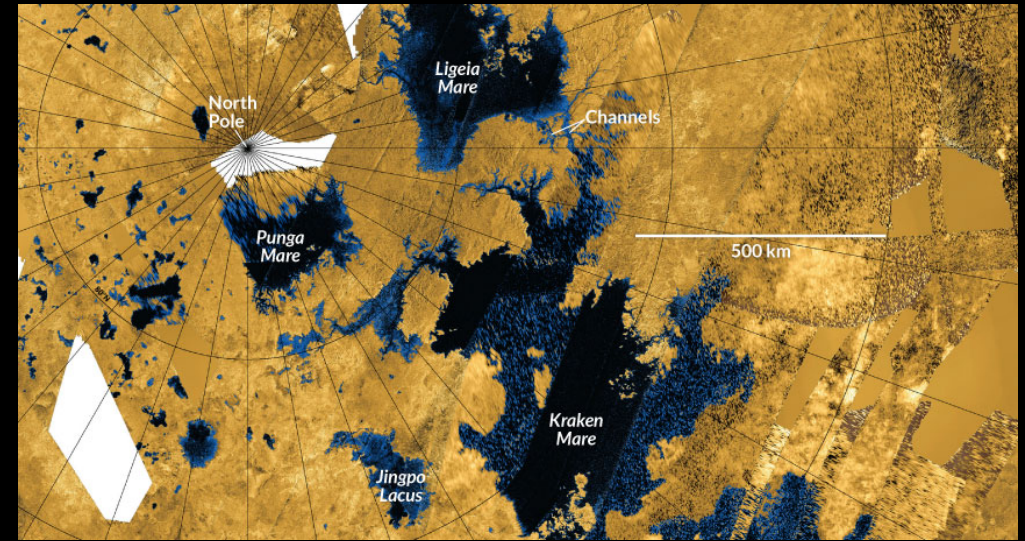
‘Downhole’ access to the foaming interface

Under-ice exploration of ocean ceiling

Open ocean exploration, including seafloor hydrothermal systems known to exist

Titan: best place to look for weird life?

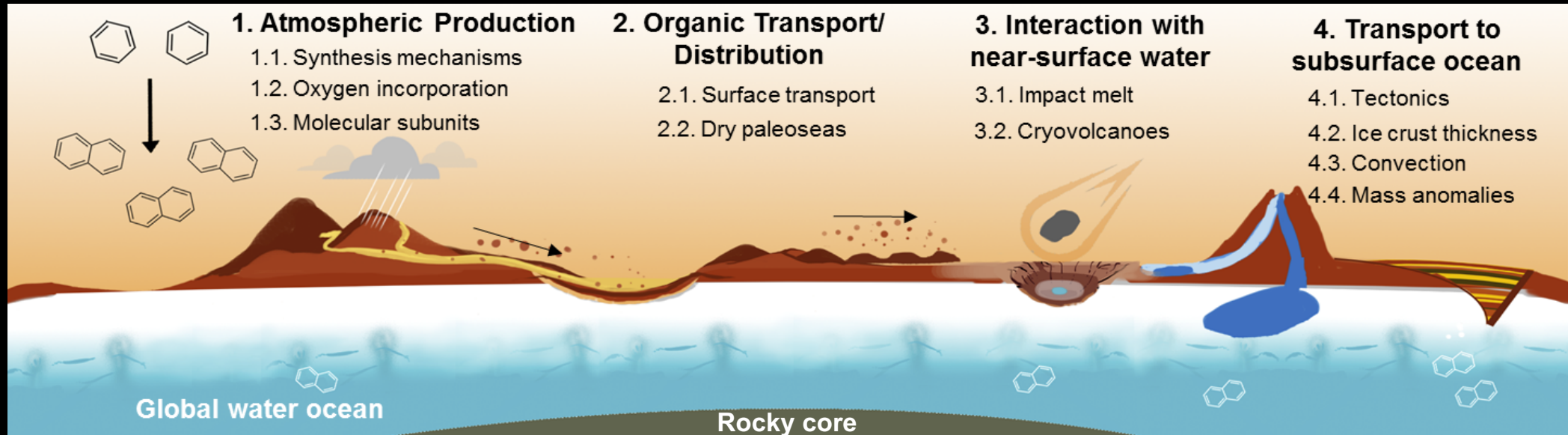
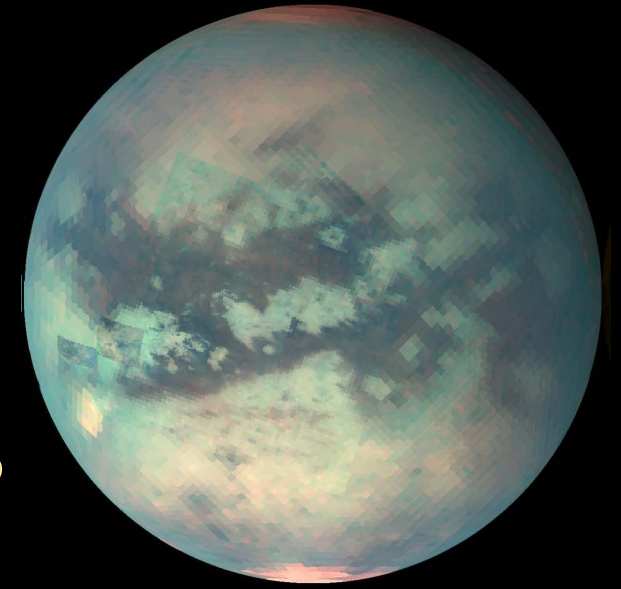
- Large organic molecules capture energy from sunlight in upper atmosphere
- Methanogenic cycle weathers organic sediment on surface
- Vast, global interior salt-water ocean
- Where and how do the organics interact with the water?



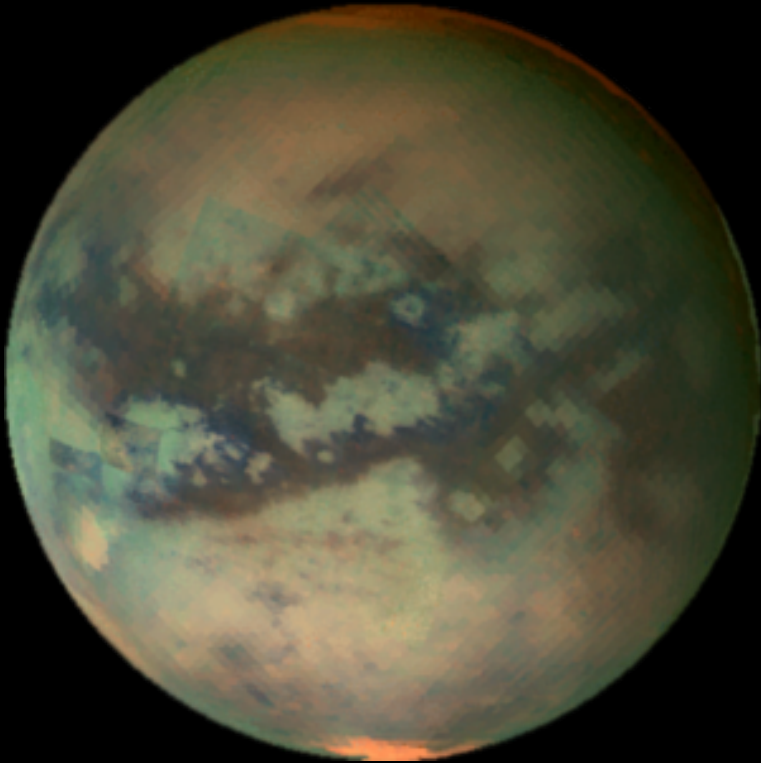
Jaumann et al., 2008

A world of two oceans

- Earth used to be a “pale orange dot” too
- How far can organic chemistry evolve without life?



A Titan program in three steps



Comprehensive reconnaissance
of a complex world

- Atmospheric organics factory
- Global surface mapping
- Gravity, tidal mapping

Aerial exploration

Buoyant sea exploration

Mobile surface exploration

In situ analysis of weathered organics

Sample return

Through-crust ocean access



From MEP to OWE

- Mars is technically moderate compared to the icy ocean worlds
- Mars missions can respond to new knowledge on half-decade cadence
- Projects managed within a program allow rapid progress
- Shared operational infrastructure 'lowers the bar' for each mission
- Core capabilities should be developed outside individual flight projects
- Medium-class directed-purpose missions form 'connective tissue'



Heavy-lift launch

- 2-yr transfer to Europa
- 3-4 years to Saturn orbit

Solar Electric Propulsion

- 25 kW-class commercial systems available today
- 5-6 years to Saturn with standard LV
- Mission-scalable



Key technologies are common to all targets

- Planetary protection of and from ocean-world material
- “Life-detection” measurement techniques and instruments
- Sample acquisition, handling, preservation
- Cryogenic mechanisms and electronics
- Modular radioisotope power systems
- Autonomous, exploratory science investigations

